Network for the Detection of Atmospheric Composition Change (NDACC)

UV-VIS Working Group



Recommendations for total ozone retrieval from NDACC zenith-sky UV-VIS spectrometers

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1. Introduction

One of the key goals of the NDACC is to ensure that long term high quality data sets of a number of important atmospheric parameters are regularly delivered and made available to the scientific community. Protocols for data acquisition, data evaluation and quality assessment are defined as part of the activities of the different working groups of the NDACC. Based on results presented and discussed at the last UV-Vis Working Group Meeting in November 2007, it was recognized that the UV-Vis total ozone data sets currently submitted to the NDACC data base still suffer from residual inconsistencies due to (1) differences in the DOAS settings applied by the different data providers, in particular as regards the source of the ozone absorption cross-sections, and (2) a lack of homogeneity in the air mass factors applied to convert O₃ slant columns into vertical columns. The aim of the present note and of its accompanying material is to provide the recommendations, tools and input data sets that are needed to improve the homogeneity of the UV-Vis total ozone measurements delivered to the NDACC data base. It is anticipated that after eventual adjustments based on feedback provided by the WG members, the guidelines provided in this document will become NDACC standards for total ozone retrievals from UV-Vis instruments, to be applied for future measurements as well as for reprocessing of historical data sets.

2. Recommended settings for O₃ vertical column retrieval

	RECOMMENDE	COMMENTS
	D SETTINGS	
Fitting interval	450-550 nm	This fitting range provides good sampling of the O ₃ differential
		absorption structures, while avoiding contamination by the
		strongest water vapor and O ₄ absorption bands.
Wavelength	Calibration based on	We recommend that measured spectra be aligned according to
calibration	reference solar atlas	Fraunhofer lines, on a atlas of solar lines such as Chance and
method		Spurr (1997) or Kurucz (1984).
Cross-sections		
O ₃	Bogumil et al,	These absorption cross-sections have been measured as part of the
	(2003), 223° K	SCIAMACHY pre-flight calibration. They are characterized by an
		excellent signal/noise ratio in the Chappuis bands and good
		consistency with values in the Huggins bands.
NO ₂	Vandaele et al.	The low temperature reference data set of Vandaele et al. (220° K)
	(1997), 220° K	is adequate for stratospheric NO ₂ retrievals and therefore adequate
		for NO_2 removal in the O_3 fitting range.
H_2O	Hitran 2004	Cross-sections generated using the latest version of the HITRAN
		data base should be used.
O ₄	Greenblatt et al.	This data set is generally accepted as the baseline O ₄ cross section
	(1990)	reference. However care must be taken to use the wavelength scale
		corrected version of Burkholder. Note that a good alternative is
		the Hermans et al. data set (unpublished, available from
		www.aeronomie.be/spectrolab/o2.htm).
Ring effect	NDACC source	For NDACC processing we recommend the use of an effective
correction	spectrum for Ring	Ring cross-section to be included in the DOAS procedure. We
method	effect correction	provide a high resolution Ring effect cross-section source
		(generated after Chance and Spurr, 1997), to be convolved at
		appropriate instrumental resolution.
Polynomial term	Polynomial of order	The polynomial filters out the broadband atmospheric attenuation
	3 to 5 maximum	due to scattering by air and particles. It may also compensate for
		broadband structures due to changing instrumental response.

The following recommendations have been jointly formulated during the November 2007 edition of the UV-VIS Working Group Meeting, held at BAS.

Intensity offset correction	Instrument dependent	The importance of this correction largely depends on the stray- light rejection capability of the spectrometer. At minimum, it is usually safe to allow for a flat offset correction.
AMF calculation	NDACC look-up	A generic look-up table of O ₃ AMFs has been generated at IASB-
	table of O ₃ AMFs	BIRA for NDACC. This table is constructed using the TOMS
		version 8 column-resolved O ₃ profile climatology. It has been
		designed to be applicable at any NDACC station (see section 3 of
		this note).
Determination of	Langley plot	If possible (depending on long-term instrumental stability) we
residual amount		recommend the use of a fixed control spectrum, selected at high
in reference		solar elevation under stable ozone conditions. The residual
spectrum		amount in this reference spectrum should be determined as
		recommended in Vaughan et al. (1997).
SZA range used	86° – 90° SZA	Best compromise between accuracy and precision is achieved in
for twilight		the 86-90° SZA range. The effective SZA of the reported O_3
averaging of		average should be explicitly mentionned in the data product.
vertical columns		

3. A new LUT of O₃ AMFs based on the TOMS version 8 ozone profile climatology

A new climatology of O_3 AMFs has been generated with the aim to homogenize and consolidate the time-series of total ozone measurements produced by UV-VIS and SAOZ spectrometers from the NDACC. As reported in the existing literature (e.g. *Sarkissian et al., 1995, 1995b, Van Roozendael et al., 1998*), differences in the radiative transfer model, pressure and temperature profiles, ozone profile shape, and wavelength can have a significant impact (up to 5-10%) on the resulting O_3 AMF values. For the sake of homogeneity it is desirable to improve the level of standardization of the UV-Vis data evaluation process. We describe a new multi-entry data base of O_3 AMFs applicable at the global scale. Its validity is tested through application at a few NDACC stations.

The proposed data base of O_3 AMFs is based on the TOMS version 8 (TV8) ozone and temperature profile climatology (*Barthia et al.*, 2004). The TV8 is a monthly-zonal climatology sorted according to the ozone column. It has been widely used for the retrieval of global total ozone fields from recent US and European UV-VIS nadir sounders (e.g. *Barthia et al.*, 2004, *Coldewey-Egbers et al.*, 2005, *Van Roozendael et al.*, 2006, *Eskes et al.*, 2005).

The parameters considered in building the look-up table (LUT) are: wavelength, ground albedo, altitude of the station, and SZA. Table 1 summarizes these different parameters and their corresponding values.

Values
- Latitude: 85°S to 85°N step 10°
- Month: 1 (Jan) to 12 (Dec) step 1
- Ozone: 125 to 575 DU step 50 DU (TOMS O ₃ grid)
440 to 580 nm step 35 nm
0 and 1
0 and 4 km
30, 50, 70, 80, 82.5, 85, 86, 87, 88, 89, 90, 91, and 92°

Table 1: Parameters of the look-up table and their corresponding values.

 O_3 AMFs have been calculated using the UVSPEC/DISORT radiative transfer (RT) model which includes a treatment of the multiple scattering in a pseudo-spherical geometry. This model has been validated through several intercomparison exercises (e.g. *Hendrick et al.*, 2006; Wagner et al., 2007). An aerosol extinction profile corresponding to background conditions has been used for the AMF calculation. It has been constructed from the aerosol model of *Shettle* (1989) included in UVSPEC/DISORT. Therefore the present O_3 AMF climatology is not suitable for volcanic conditions.

The climatology consists of 18 look-up tables (size: 1.0 MB each), each of them corresponding to one TOMS latitude (table #1: $85^{\circ}S...table$ #18: $85^{\circ}N$). An interpolation routine has been developed to extract appropriately parameterized O₃ AMFs for the different NDACC stations. Compared to version 1.0, the new version 2.0 of the routine allows AMFs to be interpolated on a yearly basis. The look-up tables are the same for both versions. The interpolation routine is written in FORTRAN 77 and a DOS executable has been created. The source code is also available for compilation on LINUX machines. In addition, a global monthly climatology of the surface albedo is coupled to the interpolation routine so that realistic albedo values can be obtained in a transparent way. This albedo climatology is extracted from the GOME surface albedo database developed by *Koelemeijer et al.* (2003). It consists of 12 look-up tables, one for each month of the year. The wavelength corresponding to these tables is 494 nm and albedo values are given for grid-cells of 1° x 1° (latitude: -89.5° to 89.5°; longitude: -179.5° to 179.5°).

4. How to use the AMF climatology?

The zip file contains 34 files: 18 O₃ AMF look-up tables, 12 surface albedo look-up tables (size: 1.1 MB each), a DOS executable ('o3_amf_interpolation_v2_0_dos'), the source code in FORTRAN 77 ('o3_amf_interpolation_v2_0.for'), and two input files for selecting parameter values. All the files should be located in the same directory for a proper use of the climatology. In the file 'input_o3_amf.dat', the user can enter a value for wavelength, latitude, longitude, surface albedo, and altitude of the station. Regarding the albedo, the user has to give a value to a flag in order to determine whether he wants to use the albedo climatology (flag=1) or not (flag=2). The user has also to define the name of the file with day numbers, SZAs and corresponding O₃ columns (here called 'Day_SZA_O3_col.dat'; maximum number of lines in this file: 500000) and to give a value to the flag for the interpolation on the O₃ column (fixed to 1 if the O₃ columns in 'Day_SZA_O3_col.dat' are vertical columns in DU and to 2 if O₃ columns are slant columns in molec/cm²). The last flag is for the display of the interpolation results on the screen (1: display; 2: no display). The resulting O₃ AMFs are stored in a file called 'o3_amf_output.dat'.

5. Verification and examples of application to NDACC data

Previous studies have demonstrated that, for AMF calculation, the UVSPEC/DISORT model shows very good consistency with others RT models (see e.g. *Hendrick et al., 2006; Wagner et al., 2007*). Nevertheless, to firmly assess the reliability of the present calculations, a quick verification exercise has been undertaken in collaboration with Dmitry Ionov (St. Petersburg State University and CNRS/SA). AMF calculations based on identical settings were performed using both UVSPEC/DISORT and the SCIATRAN model from IUP-Bremen (see www.iup.uni-bremen.de/sciatran/index.html). These show that both RT models are in excellent agreement (difference less than 2 %) when initialized in the same way.

To qualitatively illustrate the impact of using the new O_3 AMF climatology in comparison to the often used standard mid-latitude value, time-series of AMFs have been calculated for one year of data at five stations of the SAOZ/NDACC network: Sodankyla (67.4°N, 26.6°E), Jungfraujoch (46.5°N, 8°E), Observatoire de Haute Provence (OHP; 43.9°N, 5.7°E), Bauru

(22.3°S, 46°W), and Dumont d'Urville (66.7°S, 140°E). These AMFs have been compared to standard AMFs calculated with temperature and O_3 profiles extracted from the US Standard Atmosphere. The wavelength was fixed to 541 nm, the albedo to 0, and the altitude to 0 km, except for Jungfraujoch (altitude: 3.58 km). Results are presented in Figure 1.



Figure 1: Comparison between climatological (red open squares) and standard (black line) O_3 AMFs at 90° SZA at Sodankyla, Dumont d'Urville, Junfraujoch, Observatoire de Haute Provence, and Bauru.

At mid- and high-latitude in the NH, the largest difference is obtained in summer with climatological AMFs smaller than the standard AMF by up to 8%. In winter, the difference is smaller than 2% at Sodankyla and Jungfraujoch but can reach 5%

at OHP (with still climatological AMFs smaller than standard ones). At highlatitude in the SH, the climatological AMFs are smaller than the standard AMFs in summer while the opposite feature is found in winter. The relative difference is smaller than 5%. In the tropics, the climatological AMFs are systematically smaller than the standard AMFs with relative difference in the 5-10% range.

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7. Contact

For any questions, comments or bug report regarding the O_3 AMF LUT, please contact François Hendrick at the Belgian Institute for Space Aeronomy (IASB-BIRA). E-mail: franch@oma.be.

8. Acknowledgments

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9. Supplementary material

The following data files are provided together with the present document:

File name	Description
o3_amf_lut_v2_0.zip	O ₃ AMF LUT package. This contains AMF LUTs provided in 18
	latitude bands, monthly tables of global albedo values and an
	AMF extraction routine. The source code is given together with
	MS-DOS executable files.
ndacc_xsecs_v2.zip	Zip file containing reference NO ₂ , O ₃ , H ₂ O, O ₄ and Ring effect
	absorption cross-sections recommended for use in DOAS ozone
	retrievals.